

# Bottom quark contribution to spin-dependent dark matter detection

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We investigate a previously overlooked bottom quark contribution to the spin-dependent cross section for Dark Matter(DM) scattering from the nucleon. While the mechanism is relevant to any supersymmetric extension of the Standard Model, for illustrative purposes we explore the consequences within the framework of the Minimal Supersymmetric Standard Model(MSSM). We study two cases, namely those where the DM is predominantly Gaugino or Higgsino. In both cases, there is a substantial, viable region in parameter space ( $m_{\tilde{b}} - m_{\chi} \lesssim \mathcal{O}(100)$  GeV) in which the bottom contribution becomes important. We show that a relatively large contribution from the bottom quark is consistent with constraints from spin-independent DM searches, as well as some incidental model dependent constraints.

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## I. INTRODUCTION

The existence of non-baryonic Dark Matter(DM) has been established by many astronomical observations [1, 2]. Amongst the many candidates for DM, the so-called Weakly Interacting Massive Particles(WIMPs), which would have a mass in the range  $\mathcal{O}(1)$  GeV –  $\mathcal{O}(1)$  TeV, are one of the most attractive. These particles would only interact with Standard Model(SM) particles through weak interactions (and gravity), in order to yield a DM relic density consistent with measurement  $\Omega_{\text{DM}} h^2 = 0.1199 \pm 0.0027$  [3].

Direct detection of DM relies on observing the recoil energy after scattering from normal matter through weak interactions. Several DM direct detection experiments have claimed a possible excess, namely DAMA [4], Co-GeNT [5], CRESST [6] and CDMS [7]. On the other hand, these results are challenged by the absence of signals at XENON100 [8] and LUX [9], as well as CDMSlite [10] in the light DM region. The coherent, spin-independent(SI) interaction between a DM particle, generically labelled  $\chi$ , and a nucleus is proportional to the nucleon number. Because of the relatively heavy nuclei chosen for most of the above mentioned experiments, both the observed excess and stringent exclusion limits are based on SI  $\chi - p$  scattering.

As for spin-dependent(SD) DM detection [11], in a simple shell model the spin of the nucleus is that of a single, unpaired nucleon. As a consequence, the matrix element for SD  $\chi$ -nucleus scattering will be roughly comparable with that for SI  $\chi$ -nucleon scattering, with no enhancement by the nucleon number. As a result, the current DM direct searches place only very loose bounds on the SD cross section [12–14].

In the standard calculation of SD DM-nucleon scattering the heavy quark contribution is usually neglected. That is, only the contributions from  $\Delta u$ ,  $\Delta d$  and  $\Delta s$  are

included. However, as explained in the context of the proton weak charge [15], the usual decoupling of heavy quarks through the Appelquist-Carrazzone theorem [16] does not apply to quantities influenced by the U(1) axial anomaly [17–21]. In that case, rather than being suppressed by inverse powers of the heavy quark mass, the suppression is only logarithmic. These logarithmic corrections were studied in considerable detail by Bass *et al.* in Refs. [15, 22, 23], at both leading and next-to-leading order. As we shall explain here, there are interesting scenarios of supersymmetry(SUSY), generally involving a relatively light sbottom, where the logarithmic radiative correction involving the  $b$ -quark that is further enhanced by resonant effect may make a significant contribution to SD DM-nucleon scattering.

Indeed, SUSY [24, 25] is widely believed to provide the most promising explanation for new physics beyond SM. In SUSY models with R-parity conservation, the lightest supersymmetric particle(LSP) is stable and can become a DM candidate. On the other hand, both the LHC SUSY searches [26, 27] and naturalness arguments [28, 29] suggest that only the third generation supersymmetric quarks(squarks) can be light. In Ref. [30], it has been argued that an sbottom with a mass as light as  $\sim \mathcal{O}(15)$  GeV might still be consistent with current searches. In other models, such as the simplified model framework [31] and flavored DM models [32, 33], the DM can only couple to the bottom quark, as motivated by the recent DM indirect signals [34]. Studying the bottom quark contribution to the DM-nucleon SD cross section is crucial in models of this type.

In this work, we focus on the Minimal Supersymmetric Standard Model (MSSM) with a relatively light sbottom, showing when and how the bottom contribution becomes important. When the DM is Wino, there is no coupling between DM and the  $Z$ -boson and only squark mediated processes can contribute to  $\chi$ -nucleon scattering. We investigate the parameter space where the sbottom contribution is comparable to, or larger than, the first generation squark contribution. When the DM is Higgsino, the first two generation squark mediated processes

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are greatly suppressed by their small Yukawa couplings. However, the Higgsino can couple to the  $Z$ -boson. The constructive and destructive interference effects between  $Z$  and sbottom ( $\tilde{b}$ ) mediated processes are discussed in detail for a number of variations on the structure of the neutralino.

Any sbottom mediated process that contributes to the SD scattering cross section can also contribute to SI scattering. We consider the stringent LUX constraint on SI DM detection for light sbottom scenarios of interest. A relatively large SD bottom contribution can indeed be found, while maintaining consistency with the LUX constraint. We also consider several model dependent constraints from collider searches. We stress that our conclusion has implications beyond the MSSM, which is used here purely for purposes of illustration.

This paper is organised as follows. In Sec. II, we explain the theoretical framework for the calculation of the SD DM-nucleon scattering cross section. Sec. III discusses the bottom contribution for Wino and Higgsino DM. The corresponding SI detection and LHC constraints on the light sbottom scenario are considered in Sec. IV and Sec. V. In Sec. VI we present some concluding remarks.

## II. EFFECTIVE INTERACTION FOR SPIN-DEPENDENT DM-NUCLEON SCATTERING IN MSSM

Given a general effective Lagrangian

$$\mathcal{L}_{\text{SD}}^{\text{eff}} = d_q \bar{\chi} \gamma^\mu \gamma_5 \chi \bar{q} \gamma_\mu \gamma_5 q, \quad (1)$$

the spin-dependent  $\chi$ -nucleon scattering cross section can be written as

$$\sigma_{\text{SD}}^{p,n} = \frac{12}{\pi} \left( \frac{m_\chi m_{p,n}}{m_\chi + m_{p,n}} \right)^2 |a_{p,n}|^2, \quad (2)$$

where

$$a_{p,n} = \sum_q d_q \Delta_{q,p,n}. \quad (3)$$

The factors  $\Delta_{q,p,n}$  parameterise the corresponding quark spin content of the nucleon:

$$2s_\mu \Delta_{qN} = \langle N | \bar{\psi}_q \gamma_\mu \gamma_5 \psi_q | N \rangle, \quad (4)$$

where  $s_\mu$  is the nucleon spin. The preferred values of the light quark contributions in the proton and neutron are:

$$\begin{aligned} \Delta_u^{(p)} = \Delta_d^{(n)} = 0.84, \quad \Delta_d^{(p)} = \Delta_u^{(n)} = -0.43, \\ \Delta_s^{(p)} = \Delta_s^{(n)} = -0.02, \end{aligned} \quad (5)$$

where the strange quark contribution is motivated by a recent lattice QCD calculation [35].

In the MSSM at tree level, there are two processes which can contribute to the effective Lagrangian. The

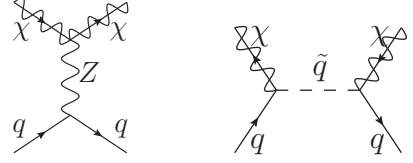


FIG. 1: Processes that contribute to the Dark Matter spin dependent cross section for scattering from a nucleon.

corresponding Feynman diagrams are given in Fig. 1. From those processes, we are able to calculate the coefficients of the effective Lagrangian from the renormalisable Lagrangian below:

$$\mathcal{L} = \bar{q}(a_q + b_q \gamma_5) \chi \bar{q} + c \bar{q} \gamma^\mu \gamma_5 q Z_\mu + d \bar{\chi} \gamma^\mu \gamma_5 \chi Z_\mu. \quad (6)$$

The corresponding couplings in the MSSM are written as [36]

$$\begin{aligned} a_u = i \frac{Z_u^L}{2} \left( \frac{-g}{\sqrt{2}c_w} \left( \frac{1}{3} Z_N^{11} s_w + Z_N^{21} c_w \right) - Y_u Z_N^{41} \right) \\ + i \frac{Z_u^R}{2} \left( \frac{2\sqrt{2}g s_w}{3c_w} Z_N^{11} - Y_u Z_N^{41} \right) \end{aligned} \quad (7)$$

$$\begin{aligned} b_u = i \frac{Z_u^L}{2} \left( \frac{g}{\sqrt{2}c_w} \left( \frac{1}{3} Z_N^{11} s_w + Z_N^{21} c_w \right) - Y_u Z_N^{41} \right) \\ + i \frac{Z_u^R}{2} \left( \frac{2\sqrt{2}g s_w}{3c_w} Z_N^{11} + Y_u Z_N^{41} \right) \end{aligned} \quad (8)$$

$$\begin{aligned} a_d = i \frac{Z_d^L}{2} \left( \frac{-g}{\sqrt{2}c_w} \left( \frac{1}{3} Z_N^{11} s_w - Z_N^{21} c_w \right) + Y_d Z_N^{31} \right) \\ + i \frac{Z_d^R}{2} \left( \frac{-\sqrt{2}g s_w}{3c_w} Z_N^{11} + Y_d Z_N^{31} \right) \end{aligned} \quad (9)$$

$$\begin{aligned} b_d = i \frac{Z_d^L}{2} \left( \frac{g}{\sqrt{2}c_w} \left( \frac{1}{3} Z_N^{11} s_w - Z_N^{21} c_w \right) + Y_d Z_N^{31} \right) \\ + i \frac{Z_d^R}{2} \left( \frac{-\sqrt{2}g s_w}{3c_w} Z_N^{11} - Y_d Z_N^{31} \right) \end{aligned} \quad (10)$$

$$c = \frac{i}{2} \frac{g}{c_w} T_{3q} \quad (11)$$

$$d = -\frac{i}{4} \frac{g}{c_w} ((Z_N^{41})^2 - (Z_N^{31})^2) \quad (12)$$

We consider first the  $Z$  boson mediated amplitude in the non-relativistic limit:

$$\begin{aligned} \mathcal{M}_{\text{SD}}^Z &= c d \bar{\chi} \gamma^\mu \gamma_5 \chi \frac{-ig_{\mu\nu}}{Q^2 - m_Z^2} \bar{q} \gamma^\nu \gamma_5 q \\ &\sim c d \frac{i}{m_Z^2} (1 + \mathcal{O}(m_Z^{-2})) \bar{\chi} \gamma^\mu \gamma_5 \chi \gamma_\mu \gamma_5 q \\ &\sim c d \frac{i}{m_Z^2} \bar{\chi} \gamma^\mu \gamma_5 \chi \gamma_\mu \gamma_5 q, \end{aligned} \quad (13)$$

so that the effective coupling  $d_q$  in Eq. (1) is:

$$d_q = \frac{cd}{m_Z^2} = \frac{g^2}{4m_W^2} T_{3q} ((Z_N^{41})^2 - (Z_N^{31})^2). \quad (14)$$

Next, for the  $\tilde{q}$  mediated process we find:

$$\begin{aligned}\mathcal{M}_{\text{SD}}^{\tilde{q}} &= \bar{\chi}(a - b\gamma_5)q \frac{i}{(p_\chi + p_q)^2 - m_{\tilde{q}}^2} \bar{q}(a + b\gamma_5)\chi \\ &\sim \frac{-i}{m_{\tilde{q}}^2 - (m_\chi + m_q)^2} \bar{\chi}(a - b\gamma_5)q \bar{q}(a + b\gamma_5)\chi \\ &= \frac{-i}{m_{\tilde{q}}^2 - (m_\chi + m_q)^2} (a^2 \bar{\chi}q \bar{q}\chi - b^2 \bar{\chi}\gamma_5 q \bar{q}\gamma_5 \chi) \\ &\ni \frac{-i}{m_{\tilde{q}}^2 - (m_\chi + m_q)^2} \left( \frac{a^2 + b^2}{4} \bar{\chi}\gamma^\mu\gamma_5\chi \bar{q}\gamma_\mu\gamma_5q \right), \quad (15)\end{aligned}$$

In this case the effective coupling in Eq. (1) is:

$$d_q = -\frac{1}{4} \frac{a^2 + b^2}{m_{\tilde{q}}^2 - (m_\chi + m_q)^2}. \quad (16)$$

Note that the tree level effective coupling  $d_q$  is only reliable when  $m_{\tilde{q}} - m_\chi$  is significantly larger than  $m_q$ . Some discussions regarding the precision of the tree level approximation are given in Appendix A. And we have also checked that the result calculated from Eq. (16) matches well with numerical tool micrOMEGAs for light flavor quark.

### III. LIGHT SBOTTOM CONTRIBUTION

For most processes of physical interest the Appelquist-Carrazzone theorem tells us that heavy quark contributions are suppressed by order  $1/m_Q^2$ . However, as explained in the introduction, because of the  $U(1)$  axial anomaly, the heavy quark contributions to spin-dependent quantities are only logarithmically suppressed. The particular case where this has been explored in great detail is the neutral weak charge of the proton. Without heavy quarks this is just  $\Delta u - \Delta d - \Delta s$ , which has been used to infer values of  $\Delta s$ . However, for a precise determination one must include the radiative corrections involving heavy quark loops which enter at order  $1/\ln m_Q$ . For example, one finds a LO correction from the  $b$ -quark equal to [15]:

$$\Delta_b^{(p)} = -\frac{6}{23\pi} \tilde{\alpha}_b (\Delta_u^{(p)} + \Delta_d^{(p)} + \Delta_s^{(p)}) \sim -0.0066. \quad (17)$$

We note that Eq. (17) is second order in the strong coupling at the  $b$  mass, as is evident in the residual 5-flavor factor  $6/23$  appearing there. However, the regularisation of the triangle diagram leads to a logarithm in  $m_b$  in the numerator which has been used to cancel the logarithm in one factor of  $\tilde{\alpha}_b$ . The logarithmic radiative correction  $\Delta_b^{(p)}$  is around 2 order of magnitude below the  $\Delta_u^{(p)}$ . We will show later that with further enhancement from resonant effect the contribution from  $\Delta_b^{(p)}$  can easily become dominant in the spin-dependent  $\chi$ -nucleon scattering.

Provided that the difference between the sbottom mass and that of the DM candidate is significantly larger than

the mass of the  $b$ -quark, the  $\tilde{q}$  propagator in Fig. 1 can be effectively factored out, leaving the familiar triangle diagram which involves the  $U(1)$  axial anomaly. In this case the bottom contribution to the axial charge of the target proton can be taken directly from Eq. (17). We shall take the running coupling  $\tilde{\alpha}_b = 0.2$ . As a result, for the  $Z$ -mediated process, the contribution of  $\Delta_b^{(p)}$  can only change the result by a factor of

$$\left( \frac{\Delta_u^{(p)} - \Delta_d^{(p)} - \Delta_s^{(p)} - \Delta_b^{(p)}}{\Delta_u^{(p)} - \Delta_d^{(p)} - \Delta_s^{(p)}} \right)^2 \sim 1.01, \quad (18)$$

which is clearly very small.

On the other hand, the term involving  $\Delta_b^{(p)}$  can give a significant contribution to the spin-dependent  $\chi$ -nucleon cross section when  $\tilde{b}$  is relatively close in mass to the DM candidate, i.e. with resonant enhancement. For simplicity we study the cases where the DM particle is either pure Wino or pure Higgsino. The corresponding couplings are:

$$a_b^{(\tilde{W})} = \frac{ig}{2\sqrt{2}} Z_b^L, \quad a_b^{(\tilde{H})} = \frac{i}{2} Y_b Z_N^{31} (Z_b^L + Z_b^R) \quad (19)$$

$$b_b^{(\tilde{W})} = -\frac{ig}{2\sqrt{2}} Z_b^L, \quad b_b^{(\tilde{H})} = \frac{i}{2} Y_b Z_N^{31} (Z_b^L - Z_b^R) \quad (20)$$

where  $Y_b = \frac{q}{\sqrt{2}m_W \cos\beta} m_b$ . So, the cross section can be written as

$$\sigma_{\text{SD}}^{\tilde{b}-\tilde{W}} = \frac{12}{\pi} \left( \frac{m_\chi m_p}{m_\chi + m_p} \right)^2 \left( -\frac{g^2 (T_{3b} Z_b^L)^2}{4(m_{\tilde{b}}^2 - (m_\chi + m_b)^2)} \Delta_b^{(p)} \right)^2 \quad (21)$$

$$\sigma_{\text{SD}}^{\tilde{b}-\tilde{H}} = \frac{12}{\pi} \left( \frac{m_\chi m_p}{m_\chi + m_p} \right)^2 \left( -\frac{0.5 Y_b^2 (Z_N^{31})^2}{4(m_{\tilde{b}}^2 - (m_\chi + m_b)^2)} \Delta_b^{(p)} \right)^2 \quad (22)$$

where we have assumed the gauge eigenstate limit and only the sbottom mediated process is contributing. By fixing  $m_\chi$  at either 10 or 100 GeV and taking  $Z_b^L = 1$  and  $\tan\beta = 40$ ,  $Z_N^{31} = \frac{1}{\sqrt{2}}$  for Wino and Higgsino DM, respectively, we can calculate the corresponding cross section as a function of  $m_{\tilde{b}}$ . The result is shown in Fig. 2. From Fig. 2 we see that the sbottom can give a very large contribution when the mass splitting  $m_{\tilde{b}} - m_\chi$  is  $\lesssim 100$  GeV.

#### A. Comparison with the contribution from the first generation squarks

First, we study the simpler case where the DM is gaugino. In this case there is no coupling between the  $Z$  boson and DM and only the squark mediated process can contribute to the SD interaction. In this subsection, we investigate the extent to which the sbottom should be lighter than first generation squark, so that they at least

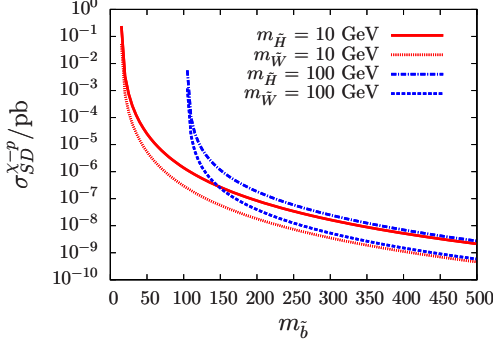


FIG. 2: Sbottom contribution to the SD scattering cross section for Wino and Higgsino Dark Matter from the proton.

have comparable cross sections. In the following we consider the sum of the contributions of all first generation squarks ( $\tilde{u}_L, \tilde{u}_R, \tilde{d}_L, \tilde{d}_R$ ), with their masses taken to be degenerate for simplicity.

Assuming that the DM is either pure Wino or Bino, the ratio of the corresponding SD cross section for sbottom to the sum of the contributions from all the first generation squarks can be calculated as

$$\frac{\sigma_{SD}^{\tilde{b}-\tilde{W}}}{\sigma_{SD}^{\tilde{q}_{u,d}-\tilde{H}}} = \frac{(Z_b^L)^2 |\Delta_b^{(p)}|}{m_b^2 - (m_\chi + m_b)^2} / \frac{\Delta_u^{(p)} + \Delta_d^{(p)}}{m_{\tilde{q}_{u,d}}^2 - m_\chi^2} \quad (23)$$

$$\frac{\sigma_{SD}^{\tilde{b}-\tilde{B}}}{\sigma_{SD}^{\tilde{q}_{u,d}-\tilde{B}}} = \frac{(a^2 + b^2)_{\tilde{b}_1}}{m_b^2 - (m_\chi + m_b)^2} \times \left( \frac{\sum_{\tilde{u}_{L,R}} (a^2 + b^2) \Delta_u^{(p)} + \sum_{\tilde{d}_{L,R}} (a^2 + b^2) \Delta_d^{(p)}}{m_{\tilde{q}_{u,d}}^2 - m_\chi^2} \right)^{-1}. \quad (24)$$

The corresponding contours of  $\sigma_{SD}^{\tilde{b}-\chi} = \sigma_{SD}^{\tilde{q}_{u,d}-\chi}$  are shown in Fig. 3. The case of Wino DM is more interesting than that of Bino DM because of its larger  $g_2$  coupling. In this case, for 1.5 TeV first generation squarks and  $\mathcal{O}(100)$  GeV DM, an sbottom lighter than about 200 GeV gives a larger cross section than the first generation squarks. On the other hand, for Bino DM, a much lighter sbottom ( $\sim 110$  GeV) is required – too light for the present calculation to be reliable.

## B. Contribution coherent with that of the Z-boson

When the DM is predominantly Higgsino, the first two generation squark mediated processes are greatly suppressed by their small Yukawa couplings. Its couplings to the Z boson and sbottom are dependent on the mixing between the two Higgsino states.

Firstly, we briefly discuss the Higgsino mixing in the MSSM. In the basis  $(\tilde{B}, \tilde{W}, \tilde{H}_d^0, \tilde{H}_u^0)$ , the neutralino mass

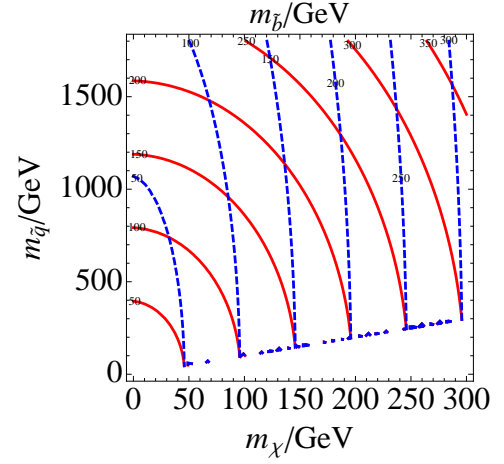


FIG. 3: Contours of constant  $m_{\tilde{b}}$  which show where the sbottom and degenerate first generation squarks give the same contribution for Wino (Red solid line) and Bino (Blue dashed line) DM.

matrix is given by:

$$M_N = \begin{pmatrix} M_1 & 0 & -c_\beta s_W m_Z & s_\beta s_W m_Z \\ 0 & M_2 & c_\beta c_W m_Z & -s_\beta c_W m_Z \\ -c_\beta s_W m_Z & c_\beta c_W m_Z & 0 & -\mu \\ s_\beta s_W m_Z & -s_\beta c_W m_Z & -\mu & 0 \end{pmatrix} \quad (25)$$

From the mass matrix we conclude that if

$$m_Z \ll \mu, M_1, M_2, \quad (26)$$

the four neutralino mass eigenstates  $\tilde{N}_i$  will be Bino  $\tilde{B}$  dominated, Wino  $\tilde{W}$  dominated and Higgsino ( $\tilde{H}_u^0 \pm \tilde{H}_d^0$ )/ $\sqrt{2}$  dominated, respectively. For example, if we also decouple the Bino and Wino from the mass matrix, the component difference between  $\tilde{H}_u^0$  and  $\tilde{H}_d^0$ , for a given mass eigenstate is

$$\Delta N_{\tilde{H}_u^0 - \tilde{H}_d^0} \propto \frac{m_Z^2}{M_i \mu} \quad (27)$$

For a few TeV gaugino and a few hundred GeV Higgsino,  $\Delta N_{\tilde{H}_u^0 - \tilde{H}_d^0} \sim \mathcal{O}(10^{-2})$ . Then, the contribution from the Z boson mediated process can be estimated by

$$\begin{aligned} \sigma_{SD}^Z &= \frac{12}{\pi} \left( \frac{m_\chi m_p}{m_\chi + m_p} \right)^2 \left( \sum_{q=u,d,s} d_q \Delta q^{(p)} \right)^2 \\ &= \frac{12}{\pi} \left( \frac{m_\chi m_p}{m_\chi + m_p} \right)^2 \left( \frac{g^2}{8m_W^2} ((Z_N^{41})^2 - (Z_N^{31})^2) \right)^2 \\ &\quad \left( \Delta_u^{(p)} \times \frac{1}{2} + \Delta_d^{(p)} \times \left(-\frac{1}{2}\right) + \Delta_s^{(p)} \times \left(-\frac{1}{2}\right) \right)^2, \end{aligned} \quad (28)$$

which is  $\sim 10^{-6}$  pb. From Fig. 2, we conclude that this corresponds to  $m_{\tilde{b}} \sim 150$  GeV for  $m_\chi \sim 100$  GeV.

To have a closer look at the coherent effects of  $Z$  boson and sbottom mediated processes, we have chosen the decoupled Wino/Bino limit, with the Higgsino DM mixing:

$$\chi = a\tilde{H}_d + b\tilde{H}_u, \quad (30)$$

where  $a^2 + b^2 = 1$  and  $b = 1.01a$ , as argued previously. This corresponds to a cross section for the  $Z$  mediated process of order  $\sim 3 \times 10^{-6}$  pb.

In this region, the  $\tilde{b}$  mediated process may also give a competitive contribution. As a result, the DM will have opposite sign coherent effects for the proton and neutron:

$$\sigma_{\text{SD}}^{\chi-(p,n)} = \frac{12}{\pi} \left( \frac{m_\chi m_p}{m_\chi + m_p} \right)^2 \left( \left( \frac{g^2((Z_N^{41})^2 - (Z_N^{31})^2)}{8m_W^2} \right) (T_{3u}\Delta_u^{(p,n)} + T_{3d}\Delta_d^{(p,n)} + T_{3s}\Delta_s^{(p,n)}) - \frac{0.5Y_b^2(Z_N^{31})^2}{4(m_b^2 - (m_\chi + m_b)^2)} \Delta_b^{(p,n)} \right)^2. \quad (31)$$

This makes the detailed consequences for SD DM scattering from real nuclei [37] potentially very complex.

We show the importance of the  $b$ -quark contribution through its coherent effects between  $Z$  mediated and sbottom mediated processes in Fig. 4. There we have

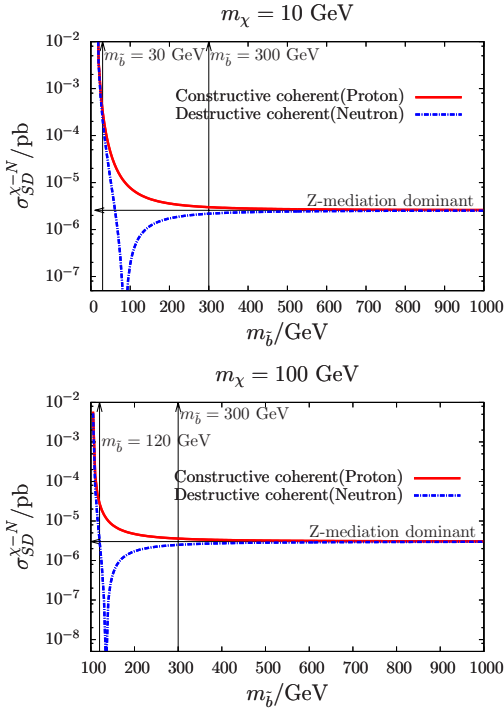


FIG. 4: Constructive and destructive coherence effects between  $Z$  mediated and  $\tilde{b}$  mediated processes for  $m_\chi = 10$  GeV (Upper) and  $m_\chi = 100$  GeV (Lower), respectively. The vertical arrowed lines at  $m_{\tilde{b}} = 30$  GeV and  $m_{\tilde{b}} = 120$  GeV indicate the value below which the calculation should not be considered reliable.

marked out the compressed spectrum region ( $\Delta(m_{\tilde{b}} - m_\chi) \lesssim 20$  GeV) where our tree level calculation cannot

be considered reliable. According to the Eq. (31) with assumed  $Z_N^{41} = 1.01Z_N^{31}$ , the first term in the parenthesis is positive for proton and negative for neutron, while the second term is always negative because  $\Delta_b^{(p)} = \Delta_b^{(n)} < 0$ . As a result, the interference terms for proton and neutron are constructive and destructive, respectively. We find that for the DM mass around  $\mathcal{O}(10 - 100)$  GeV an sbottom with mass  $\lesssim 300$  GeV can make a non-negligible contribution. In some specific regions, the corresponding cross section for the  $Z$ -mediated process may even be enhanced or reduced by several orders of magnitude.

#### IV. SPIN-INDEPENDENT DM DETECTION CONSTRAINT FROM LUX

The same process shown in the right panel of Fig. 1, which can give rise to an enhancement of the spin-dependent scattering cross section, can also contribute to spin-independent scattering. As a result, the very stringent spin-independent DM search bound from LUX [9] may already exclude some of the parameter region found to be of interest here.

We start with the following effective Lagrangian [38–41]:

$$\mathcal{L}_{\text{SI}} = \sum_q (f_q m_q \bar{\chi} \chi \bar{q} q + \frac{g_q^{(1)}}{m_\chi} \bar{\chi} i \partial^\mu \gamma^\nu \chi \mathcal{O}_{\mu\nu}^q + \frac{g_q^{(2)}}{m_\chi^2} \bar{\chi} (i \partial^\mu) (i \partial^\nu) \chi \mathcal{O}_{\mu\nu}^q) + f_G \bar{\chi} \chi G_{\mu\nu}^a G^{a\mu\nu} \quad (32)$$

where  $\chi$  is the DM field,  $m_\chi$  its mass and the twist-2 operator:

$$\mathcal{O}_{\mu\nu}^q = \frac{1}{2} \bar{q} i (D_\mu \gamma_\nu + D_\nu \gamma_\mu - \frac{1}{2} g_{\mu\nu} \not{D}) q. \quad (33)$$

The corresponding spin-independent scattering cross section of DM with a proton can be written as

$$\sigma_{\text{SI}}^{\chi-p} = \frac{4}{\pi} \mu^2 (f_N)^2 \quad (34)$$

where  $\mu = m_\chi m_N / (m_\chi + m_N)$  and

$$\frac{f_N}{m_N} = \sum_{q=u,d,s} f_q f_{T_q} + \sum_{q=u,d,s,c,b} \frac{3}{4} (q(2) + \bar{q}(2)) (g_q^{(1)} + g_q^{(2)}) - \frac{8\pi}{9\alpha_s} f_{T_G} f_G \quad (35)$$

$$\sim \sum_{q=u,d,s} f_q f_{T_q} + \sum_{q=u,d,s,c,b} \frac{3}{4} (q(2) + \bar{q}(2)) (g_q^{(1)} + g_q^{(2)}) + \frac{2}{27} \sum_{Q=c,b,t} f_{T_Q} f_Q \quad (36)$$

The light quark parameters  $f_{T_q}$  are defined by

$$f_{T_q} m_N = \langle N | m_q \bar{q} q | N \rangle, \quad (37)$$



and  $f_{T_G} = 1 - \sum_{q=u,d,s} f_{T_q}$ . Recent lattice simulations give [42–45]:

$$f_u^p = 0.023, \quad f_d^p = 0.033, \quad f_s^p = 0.026. \quad (38)$$

The second moments of the parton distribution functions (PDFs) can be used to evaluate the matrix element of  $\mathcal{O}_{\mu\nu}^q$ :

$$(p_\mu p_\nu - \frac{1}{4} m_N^2 g_{\mu\nu})(q(2) + \bar{q}(2)) = m_N < N(p) | \mathcal{O}_{\mu\nu}^q | N(p) >, \quad (39)$$

which from the CTEQ PDFs [46] yields

$$b(2) = 0.012, \quad \bar{b}(2) = 0.012, \quad (40)$$

at the  $Z$  boson mass scale.

Using a similar technique to that used in calculating the SD effective coefficient,  $d_q$ , above, we can find the corresponding effective coefficient for the spin-independent case. Based on the renormalizable Lagrangian Eq. (6), we have

$$f_q = \frac{m_\chi}{(m_{\tilde{q}}^2 - (m_\chi + m_q)^2)^2} \frac{a_q^2 + b_q^2}{8} - \frac{1}{m_q(m_{\tilde{q}}^2 - (m_\chi + m_q)^2)} \frac{a_q^2 - b_q^2}{4} \quad (41)$$

$$g_q^{(1)} + g_q^{(2)} = \frac{m_\chi}{(m_{\tilde{q}}^2 - (m_\chi + m_q)^2)^2} \frac{a_q^2 + b_q^2}{2}. \quad (42)$$

As a result the twist-2 operator,  $\mathcal{O}_{\mu\nu}^q$ , gives a much larger contribution than  $f_q$  in most cases. For Higgsino DM, where  $a_q \neq b_q$ , the second term of  $f_q$  can easily become dominant. However in this case it is negative, so a cancellation between  $f_q$  and  $g_q$  may happen in some of the parameter regions.

We first consider the pure Wino DM case, with only  $\tilde{L}$  mediated scattering. From Eq. (36), we have

$$f_N = m_p \left( \frac{3}{4} (b(2) + \bar{b}(2)) (g_b^{\tilde{L}-\tilde{W}}) + \frac{2}{27} f_{T_G} f_b^{\tilde{L}-\tilde{W}} \right) \quad (43)$$

where

$$f_b^{\tilde{L}-\tilde{W}} = \frac{g^2 m_\chi}{32} \frac{1}{(m_b^2 - m_\chi^2)^2} \quad (44)$$

$$g_b^{\tilde{L}-\tilde{W}} = \frac{g^2 m_\chi}{8} \frac{1}{(m_b^2 - m_\chi^2)^2} \quad (45)$$

The tree level calculation for the spin-independent and spin-dependent cross sections is shown in Fig. 5. We conclude from the figure that  $m_{\tilde{b}} - m_\chi \gtrsim 50$  GeV is required to evade the spin-independent bound from LUX for Wino DM. It has to be noted that the pole at  $m_{\tilde{b}} = m_b + m_\chi$ , for SI tree level results, will not show up when the full NLO effects are taken into account [47]. We have checked that our results fit the numerical results from micrOMEGAs [48, 49] quite well, outside the pole region.

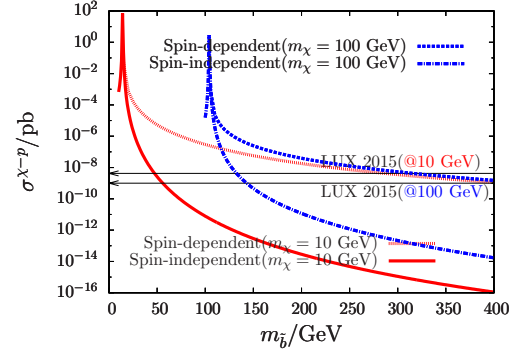


FIG. 5: SD and SI scattering cross section for Wino DM from the proton, with  $m_\chi = 10$  GeV and  $m_\chi = 100$  GeV, respectively.

Next, we discuss the more interesting case where the DM is predominantly Higgsino. As discussed above, in this case the relatively large spin dependent cross section from the sbottom mediated process can interfere coherently with the  $Z$  mediated process, leading to very different SD scattering rates for protons and neutrons.

The SI DM-proton effective coupling is

$$f_N = m_p \left( \frac{3}{4} (b(2) + \bar{b}(2)) (g_b^{\tilde{b}_1-\tilde{H}}) + \frac{2}{27} f_{T_G} f_b^{\tilde{b}_1-\tilde{H}} \right) \quad (46)$$

where

$$f_b^{\tilde{b}_1-\tilde{H}} = \frac{m_\chi}{(m_{\tilde{b}}^2 - (m_\chi + m_b)^2)^2} \frac{0.5 Y_b^2 (Z_N^{31})^2}{8} - \frac{1}{m_b(m_{\tilde{b}}^2 - (m_\chi + m_b)^2)} \frac{Y_b^2 (Z_N^{31})^2 Z_b^L Z_b^R}{4} \quad (47)$$

$$g_b^{\tilde{b}_1-\tilde{H}} = \frac{m_\chi}{(m_{\tilde{b}}^2 - (m_\chi + m_b)^2)^2} \frac{0.5 Y_b^2 (Z_N^{31})^2}{2}. \quad (48)$$

The corresponding tree level calculation for the spin-dependent and spin-independent cross sections is shown in Fig. 6. From that figure we see that a small component of left-handed sbottom is favoured in order to evade the LUX bound. When the left-handed sbottom component is relatively large, a large SD cross section may also be consistent with the LUX experiment if there is a cancellation in  $\sigma_{\text{SI}}$ .

## V. MODEL DEPENDENT CONSTRAINTS AND A GENERAL ARGUMENT

We have presented a representative study of the potential importance of the bottom quark contribution to DM spin-dependent detection within the framework of the MSSM. This particular contribution has hitherto been overlooked. However, in a realistic model such as MSSM, there are many other incidental constraints. We will briefly outline how these may be evaded, while keeping our discussion as general as possible in this section.

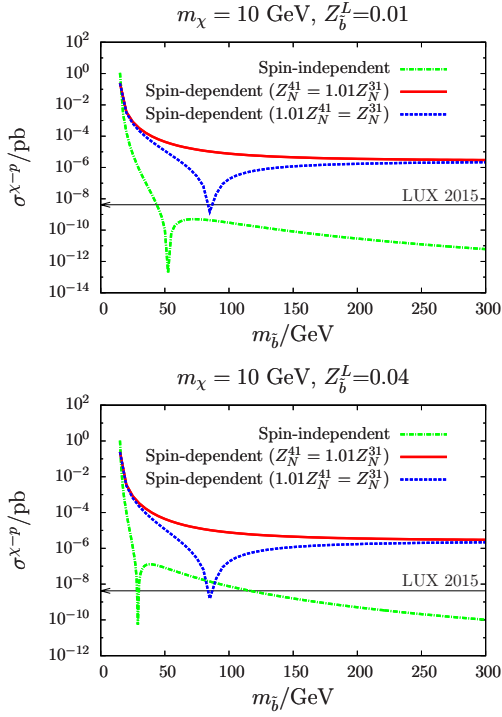


FIG. 6: SD and SI scattering cross section for different mixing of Higgsino DM from the proton. Upper:  $m_\chi = 10$  GeV,  $Z_b^L = 0.01$ . Lower:  $m_\chi = 10$  GeV,  $Z_b^L = 0.04$ .

LEP placed a very stringent bound on the chargino mass ( $m_{\tilde{H}^\pm(\tilde{W}^\pm)} > 92.4(91.9)$  GeV) [50]. Because for either Wino and Higgsino DM there is a charged partner (chargino), which has very similar mass with the DM, we cannot have Wino and Higgsino DM of  $m_\chi \lesssim 90$  GeV in a typical MSSM framework.

As for  $m_\chi \gtrsim 100$  GeV, on the other hand, it will be constrained by LHC sbottom searches [51, 52] and mono-jet search [53], since we usually need a relatively light sbottom to enhance the bottom quark spin dependent contribution. The corresponding LHC exclusion bounds and spin-dependent  $\chi - p$  scattering cross section in the  $m_{\tilde{b}} - m_\chi$  plane are shown in Fig. 7. To generate this figure we have used Eq. (21) and (22), where only the sbottom mediated process is considered. The contours of  $\sigma_{SD}^{\chi-p}$  show the condition when the sbottom mediated contribution is half the size of Z mediated process for Wino DM and a typical Higgsino DM candidate with  $Z_N^{41} = 1.01 Z_N^{31}$ . This figure suggests that a large portion of parameter space is excluded by the LHC sbottom searches.

However, there are several ways to avoid these constraints:

- For  $m_\chi \lesssim 100$  GeV, we can work in a more general framework, where the dark matter does not have any charged partners. Its couplings to the Z boson and  $\tilde{b}$  may be of the same order; e.g. the simpli-

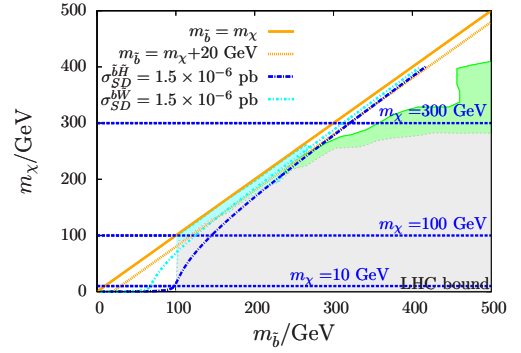


FIG. 7: LHC exclusion bound on sbottom mass versus DM mass. The grey, green and cyan shaded region correspond to the exclusion limits given by LHC searches for sbottom at 8 TeV [51], 13 TeV [51] and searches for mono-jet at 8 TeV [53], respectively.

fied model framework [31] or flavored dark matter models [32, 33].

- For  $m_\chi \lesssim 100$  GeV, if the charged Higgsino decays into DM and a relatively long lived particle, with lifetime of  $\mathcal{O}(10-100)$  cm, similar to Ref. [54]. As a result, the reconstructed track will not point to the interaction point and would therefore be unlikely to be considered a "good" track. In this case, the LEP constraints on charginos can be evaded. The light sbottom constraint can also be evaded by tuning appropriate mixing – see e.g. Refs. [30, 55].
- For  $m_\chi \gtrsim 100$  GeV, if the sbottom is decayed in more complicate modes other than  $\tilde{b} \rightarrow b\chi$ , the corresponding LHC bound on sbottom mass can be loosened.
- We can also work with heavier DM, e.g.  $m_\chi = 300$  GeV, for example, as shown in Fig. 8. In this case,  $m_{\tilde{b}} \lesssim 350$  GeV is consistent with the LHC searches, while the sbottom mediated process can give a significant contribution to spin-dependent scattering.

## VI. CONCLUSION

In this work, we have demonstrated the potential importance of the bottom quark contribution to the DM spin-dependent cross section due to the axial anomaly and resonant enhancement, which has hitherto been overlooked. Even though our calculation was carried out within the framework of the MSSM, the general conclusion will be relevant to any models with similar particle content, since the only relevant ingredients are  $\chi$ ,  $\tilde{b}$  and the Z-boson.

In the MSSM, we calculated the bottom quark contribution to spin-dependent  $\chi - p$  scattering. Firstly, we

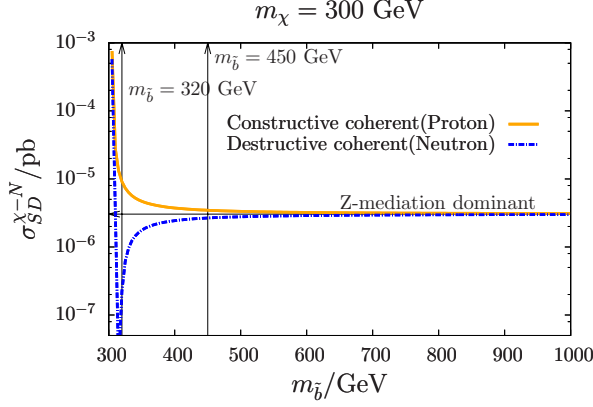


FIG. 8: Constructive and destructive interference effects between  $Z$  mediated and  $\tilde{b}$  mediated processes for Higgsino DM with  $m_\chi = 300 \text{ GeV}$ .

considered Gaugino DM, where there is no coupling between the  $Z$ -boson and DM. Assuming  $m_\chi = 100 \text{ GeV}$  and degenerate first generation squarks with a mass of  $1.5 \text{ TeV}$ , we found that an sbottom of mass  $m_{\tilde{b}} \lesssim 200 \text{ GeV}$  can give rise to a larger spin-dependent cross section for Wino DM. By contrast, for Bino DM a much lighter sbottom mass ( $m_{\tilde{b}} \sim 110 \text{ GeV}$ ) is required to give a competitive cross section. As for Higgsino DM, the first generation squark contributions are suppressed by their small Yukawa couplings. However, the  $Z$ -boson mediated process does contribute. For a given Higgsino mixing of DM the sbottom mediated process may interfere either constructively or destructively with  $Z$ -boson mediated processes, with different signs for protons and neutrons. For a typical mixing of Higgsino DM with mass around  $\mathcal{O}(10 - 100) \text{ GeV}$ , we find that an sbottom of mass below  $300 \text{ GeV}$  can have non-negligible effects.

The squark mediated process that gives rise to an increase in spin-dependent DM scattering can also contribute to the spin-independent cross section. Our calculation shows that  $\Delta(m_{\tilde{b}} - m_\chi) \gtrsim 50 \text{ GeV}$  is required to evade the LUX constraint for Wino DM, while for Higgsino DM, either a small component of left-handed sbottom or a large cancellation in  $\sigma_{SI}$  is needed. Some incidental model dependent constraints from LEP and the LHC are considered as well. Those constraints, however, can be evaded in more general theoretical frameworks.

As pointed out earlier, our tree level results may break down as the sbottom and DM masses become degenerate ( $\Delta(m_{\tilde{b}} - m_\chi) \lesssim 20 \text{ GeV}$ ). We leave the higher order calculation for this small region for future work. Finally, while the calculations for the top quark case will be more complicated because there is no clear separation of mass scales for interesting ranges of DM mass, there is a clear need to investigate the role of the axial anomaly for that case too.

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## Appendix A: Precision of tree level approximation

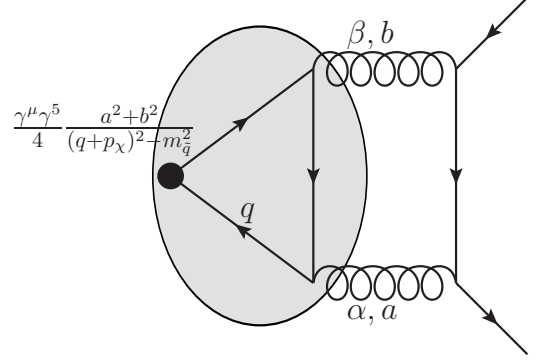


FIG. 9: Bottom quark contribution to the axial current.

The heavy quark contributions to the axial charge start at two loop level through the process shown in Fig. 9. A detailed calculation of this diagram is given in Ref. [22, 23]. In this study, all we need to know is the  $m_{\tilde{q}}$  dependence of the amplitude. Then we can derive the range of  $m_{\tilde{q}}$  for which the tree level effective coupling Eq. 16 is justified.

The  $m_{\tilde{q}}$  dependence only exists in the triangle loop that is marked by the grey shaded ellipse in Fig. 9. The vertex amplitude is

$$\Gamma_{\mu\alpha\beta}^{ab} = \frac{a^2 + b^2}{4} (-1) g^2 \int \frac{d^4 q}{(2\pi)^4} \text{Tr} \left\{ \frac{\gamma_\mu \gamma_5}{(q + p_\chi)^2 - m_{\tilde{q}}^2} \cdot \frac{i}{\not{q} - m} (i\gamma_\alpha \frac{1}{2} \lambda^a) \frac{i}{\not{q} - m} (i\gamma_\beta \frac{1}{2} \lambda^b) \frac{i}{\not{q} - m} \right\}, \quad (\text{A1})$$

where  $m, q$  is the quark mass and momentum in the triangle,  $p_\chi$  is the dark matter four momentum and  $m_{\tilde{q}}$  is the squark mass. After introducing Feynman parameter  $x$  and the substitution:

$$l^\mu = q^\mu + x p_\chi^\mu, \quad (\text{A2})$$

$$\Delta = x^2 p_\chi^2 - x p_\chi^2 + m^2(1 - x) + x m_{\tilde{q}}^2, \quad (\text{A3})$$

the vertex amplitude Eq. A1 can be simplified to

$$\Gamma_{\mu\alpha\beta}^{ab} \propto 3! \epsilon_{\alpha\beta\mu\rho} \int_0^1 dx \left\{ -2x(p_\chi)_\sigma \int \frac{d^4 l}{(2\pi)^4} \frac{l^\rho l^\sigma}{(l^2 - \Delta)^4} - x p_\chi^\rho \cdot \int \frac{d^4 l}{(2\pi)^4} \frac{l^2}{(l^2 - \Delta)^4} - x p_\chi^\rho (x^2 p_\chi^2 - m^2) \int \frac{d^4 l}{(2\pi)^4} \frac{1}{(l^2 - \Delta)^4} \right\}. \quad (\text{A4})$$



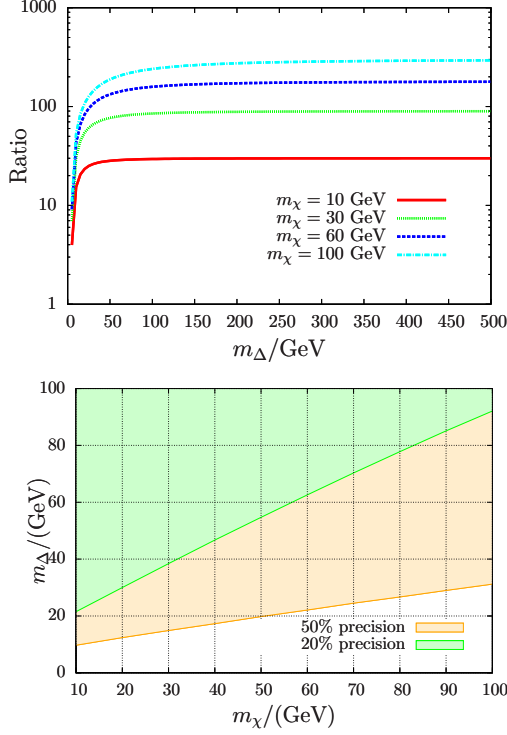


FIG. 10: Upper: The  $m_\Delta$  dependence of the ratio for several given DM masses. Lower: In the shaded region, the tree level approximation matches the loop level results within required precision.

Finally, after integrating out the  $l^\mu$ , we will get a simple  $m_{\bar{q}}$  and  $p_\chi$  dependence of  $\Gamma_{\mu\alpha\beta}^{ab}$ :

$$\Gamma^{\text{loop}}(m_{\bar{q}}, p_\chi) = \int_0^1 dx \left( \frac{3xp_\chi^\rho}{\Delta} - \frac{xp_\chi^\rho(x^2p_\chi^2 - m^2)}{\Delta^2} \right) \quad (\text{A5})$$

On the other hand, the  $m_{\bar{q}}$  dependence of the tree level effective coupling can be factored out as

$$\Gamma^{\text{tree}} \propto \frac{1}{(m^2 - m_\chi^2)^2 - m_{\bar{q}}^2}. \quad (\text{A6})$$

So we can define the ratio

$$\begin{aligned} \text{Ratio} &\propto \Gamma^{\text{loop}}/\Gamma^{\text{tree}} \\ &= ((m^2 - m_\chi^2)^2 - m_{\bar{q}}^2) \cdot \Gamma^{\text{loop}}(m_{\bar{q}}, p_\chi). \end{aligned} \quad (\text{A7})$$

Taking the non-relativistic limit for the DM momentum, i.e.  $p_\chi = (m_\chi, 0, 0, 0)$ ,  $m = m_b$  and  $m_{\bar{q}} = m_\Delta + m_\chi$ , we solve the ratio numerically. The results are shown in the upper panel of Fig. 10. From the figure we find that the ratio tends to a constant in the heavy squark region for given  $m_\chi$ , which means the tree level description is accurate. While in the region of small mass splitting, the tree level results deviate from the full loop calculation considerably by a amount.

The range of  $m_\Delta$  that permits the tree level approximation in required precision  $P$  can be solved by using the inequality

$$\text{Ratio}(m_\Delta)/\text{Ratio}(m_\Delta = 500) > (1 - P) \quad (\text{A8})$$

at each given DM mass. In the lower panel of Fig. 10, we show the  $m_\Delta$  region in which the tree level approximation matches the loop level result within 20% and 50% precision, respectively. For example, when  $m_\chi \sim 10$  GeV,  $m_\Delta \gtrsim 20$  GeV is sufficient to guarantee that the tree level approximation is accurate within 20%.

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